# Data intensive electromagnetic scattering simulation for planetary radar data analysis

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### **MARSIS** instrument

Mars Advanced Radar for Subsurface and Lonosphere Sounding <u>To map the distribution of water in the upper portions of the crust of Mars</u>

Synthetic-aperture orbital sounding radar
On board the ESA spacecraft *Mars Express*Two 20 meters elements nadir-looking dipole antenna
Centre frequencies: I.8, 3.0, 4.0, 5.0 MHz (I.0 MHz bandwidth) chirp

Radar echoes will contain both surface and subsurface reflection components
Surface backscattering from off-nadir directions is called "clutter"
Clutter can mask, or be mistaken for, subsurface echoes

# **MARSIS clutter simulation**

Facet model of rough surface scattering + 500-meter resolution elevation map of Mars from NASA MOLA instrument (*Mars Global Surveyor*)

Surface echo simulation Nouvel et al. (Radio Science vol.39, 2004)



The facet model of rough surface scattering Rees, W. G. (1990), *Physical Principles of Remote Sensing* 



Facet definition Nouvel et al. (*Radio Science vol.39, 2004*)

## **MARSIS** clutter simulation



SHARAD observation (orbit 1948)

#### Clutter simulation

MOLA topography

SHARAD (SHAllow RADar) - carried by NASA Mars Reconnaissance Orbiter

# **MARSIS clutter simulator**



# Parallelization



# **Scalability**

#### Tested on CINECA PLX and Fermi BG/Q clusters



# **Required resources**

Time required for the calculation of a testing orbit (#2665):

PLX (336 cores, 10%): 1.7 hours

Fermi (4096 cores, 2.5%): 0.43 hours (all the orbits: 180 days)

Input (orbit data, MARSIS parameters, Mars elevation): ~ 5GB Output (simulated echoes): ~ 600GB

Memory allocation: I.2TB (per orbit, 4096 cores on Fermi)

•Parallelization allows calculation on suitable hardware

- Computing resources allocation requested in 7th PRACE call
  Since the simulations of each echo and each orbit are independent, MPI is not necessary
- •Application can be suitable for a Map-Reduce approach